

5

Matter Waves

$$5-1 \quad \lambda = \frac{h}{p} = \frac{h}{mv} = \frac{6.63 \times 10^{-34} \text{ Js}}{1.67 \times 10^{-27} \text{ kg}} (10^6 \text{ m/s}) = 3.97 \times 10^{-13} \text{ m}$$

$$5-3 \quad \lambda = \frac{h}{p} = \frac{h}{mv} = \frac{6.63 \times 10^{-34} \text{ Js}}{74 \text{ kg}} (5 \text{ m/s}) = 1.79 \times 10^{-36} \text{ m}$$

$$5-5 \quad (a) \quad \lambda = \frac{h}{p} \text{ or } p = \frac{h}{\lambda} = \frac{hc}{\lambda c} = \frac{1240 \text{ eV nm}}{(10 \text{ nm})(c)} = \frac{124 \text{ eV}}{c}. \text{ As}$$

$$K = E - mc^2 = [p^2 c^2 + (mc^2)^2]^{1/2} - mc^2,$$

we must use the relativistic expression for K , when $pc \approx mc^2$. Here

$pc = 124 \text{ eV} \ll mc^2 = 0.511 \text{ MeV}$, so we can use the classical expression for K , $K = \frac{p^2}{2m}$.

$$K = \frac{p^2}{2m} = \frac{p^2 c^2}{2mc^2} = \frac{(124 \text{ eV})^2}{2(0.511 \text{ MeV})} = 0.150 \text{ eV}$$

(b) Electrons with $\lambda = 0.10 \text{ nm}$ $p = \frac{hc}{\lambda c} = \frac{12400 \text{ eV}}{c}$ as in (a). As $pc \ll mc^2 = 0.511 \text{ MeV}$, use

$$K = \frac{p^2}{2m} = p^2 c^2 = \frac{(12400)^2 (\text{eV})^2}{(2)(0.511 \times 10^6 \text{ eV})} = 150 \text{ eV}.$$

(c) Electrons with $\lambda = 10 \text{ fm} = 10 \times 10^{-15} \text{ m}$, $p = \frac{hc}{\lambda c} = \frac{1.24 \times 10^3 \text{ MeV}}{c}$. As

$pc \gg mc^2 = 0.511 \text{ MeV}$, use

$$K = [p^2 c^2 + (mc^2)^2]^{1/2} - mc^2 = pc - mc^2 = 1240 \text{ MeV} - 0.511 \text{ MeV} = 1239 \text{ MeV}.$$

For alphas with $mc^2 = 3726 \text{ MeV}$:

(a) p still is $\frac{124 \text{ eV}}{c}$. As $pc \ll 3726 \text{ MeV}$, we use $K = \frac{p^2}{2m}$:

$$K = \frac{p^2 c^2}{2mc^2} = \frac{(124 \text{ eV})^2}{(2)(3726 \text{ MeV})} = 2.06 \times 10^{-6} \text{ eV}.$$

(b) For alphas with $\lambda = 0.10 \text{ nm}$, $p = \frac{12400 \text{ eV}}{c}$. As $pc \ll mc^2 = 3726 \text{ MeV}$,

$$K = \frac{p^2}{2m} = \frac{p^2 c^2}{2mc^2} = \frac{(12400 \text{ eV})^2}{(2)(3726 \text{ MeV})} = 0.0206 \text{ eV}.$$

(c) $p = \frac{1.24 \times 10^3 \text{ MeV}}{c}$ and $pc = 1240 \text{ MeV} \sim mc^2 = 3726 \text{ MeV}$. We use

$$K = \left[p^2 c^2 + (mc^2)^2 \right]^{1/2} - mc^2 = \left[(1240 \text{ MeV})^2 + (3726 \text{ MeV})^2 \right]^{1/2} - 3726 \text{ MeV} \\ = 201 \text{ MeV}.$$

5-7 A 10 MeV proton has $K = 10 \text{ MeV} \ll 2mc^2 = 1877 \text{ MeV}$ so we can use the classical expression $p = (2mK)^{1/2}$. (See Problem 5-2)

$$\lambda = \frac{h}{p} = \frac{hc}{[(2)(938.3 \text{ MeV})(10 \text{ MeV})]^{1/2}} = \frac{1240 \text{ MeV fm}}{[(2)(938.3)(10)(\text{MeV})^2]^{1/2}} = 9.05 \text{ fm} = 9.05 \times 10^{-15} \text{ m}$$

5-9 $m = 0.20 \text{ kg}$: $mgh = \frac{mv^2}{2}$: $v = (2gh)^{1/2}$

$$p = mv = m(2gh)^{1/2} = (0.20)[2(9.80)(50)]^{1/2} = 6.261 \text{ kg} \cdot \text{m/s}$$

$$\lambda = \frac{h}{p} = \frac{6.626 \times 10^{-34} \text{ J} \cdot \text{s}}{6.261 \text{ kg} \cdot \text{m/s}} = 1.06 \times 10^{-34} \text{ m}$$

5-11 (a) In this problem, the electron must be treated relativistically because we must use relativity when $pc \approx mc^2$. (See problem 5-5). the momentum of the electron is

$$p = \frac{h}{\lambda} = \frac{6.626 \times 10^{-34} \text{ J} \cdot \text{s}}{10^{-14} \text{ m}} = 6.626 \times 10^{-20} \text{ kg} \cdot \text{m/s}$$

and $pc = 124 \text{ MeV} \gg mc^2 = 0.511 \text{ MeV}$. The energy of the electron is

$$E = (p^2 c^2 + m^2 c^4)^{1/2} \\ = \left[(6.626 \times 10^{-20} \text{ kg} \cdot \text{m/s})^2 (3 \times 10^8 \text{ m/s})^2 + (0.511 \times 10^6 \text{ eV})^2 (1.602 \times 10^{-19} \text{ J/eV})^2 \right]^{1/2} \\ = 1.99 \times 10^{-11} \text{ J} = 1.24 \times 10^8 \text{ eV}$$

so that $K = E - mc^2 \approx 124 \text{ MeV}$.

- (b) The kinetic energy is too large to expect that the electron could be confined to a region the size of the nucleus.

5-13 A canceling wave will be produced when the path length difference between the surface reflection and the reflection from the n th plane below the surface equals some whole number of wavelengths plus $\frac{\lambda}{2}$. As the path length difference between a surface reflection and a reflection from plane n is given by $(n)(1.01\lambda)$, we find that a reflection from the 50th plane has a path difference of 50.5λ with the surface reflection, and consequently cancels the surface reflection. Essentially all waves reflected at θ will cancel as the wave reflected from the second plane will be cancelled by a reflection from the 51st plane and so on.

5-15 For a free, non-relativistic electron $E = \frac{m_e v_0^2}{2} = \frac{p^2}{2m_e}$. As the wavenumber and angular frequency of the electron's de Broglie wave are given by $p = \hbar k$ and $E = \hbar \omega$, substituting these results gives the dispersion relation $\omega = \frac{\hbar k^2}{2m_e}$. So $v_g = \frac{d\omega}{dk} = \frac{\hbar k}{m_e} = \frac{p}{m_e} = v_0$.

5-17 $E^2 = p^2 c^2 + (m_e c^2)^2$
 $E = [p^2 c^2 + (m_e c^2)^2]^{1/2}$. As $E = \hbar \omega$ and $p = \hbar k$

$$\hbar \omega = [\hbar^2 k^2 c^2 + (m_e c^2)^2]^{1/2} \text{ or}$$

$$\omega(k) = \left[k^2 c^2 + \frac{(m_e c^2)^2}{\hbar^2} \right]^{1/2}$$

$$v_p = \frac{\omega}{k} = \frac{[k^2 c^2 + (m_e c^2 / \hbar)^2]^{1/2}}{k} = \left[c^2 + \left(\frac{m_e c^2}{\hbar k} \right)^2 \right]^{1/2}$$

$$v_g = \left. \frac{d\omega}{dk} \right|_{k_0} = \frac{1}{2} \left[k^2 c^2 + \left(\frac{m_e c^2}{\hbar} \right)^2 \right]^{-1/2} 2kc^2 = \frac{kc^2}{[k^2 c^2 + (m_e c^2 / \hbar)^2]^{1/2}}$$

$$v_p v_g = \left\{ \frac{[k^2 c^2 + (m_e c^2 / \hbar)^2]^{1/2}}{k} \right\} \left\{ [k^2 c^2 + (m_e c^2 / \hbar)^2]^{1/2} \right\} = c^2$$

Therefore, $v_g < c$ if $v_p > c$.

5-19 $K = \frac{mv^2}{2} = \frac{p^2}{2m}$; $(1 \times 10^6 \text{ eV})(1.6 \times 10^{-19} \text{ J/eV}) = \frac{p^2}{2(1.67 \times 10^{-27} \text{ kg})} \Rightarrow p = 2.312 \times 10^{-20} \text{ kg} \cdot \text{m/s}$,

$$\Delta p = 0.05p = 1.160 \times 10^{-21} \text{ kg} \cdot \text{m/s} \text{ and } \Delta x \Delta p = \frac{\hbar}{4\pi}. \text{ Thus}$$

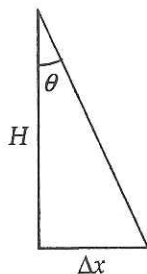
$$\Delta x = \frac{6.63 \times 10^{-34} \text{ J} \cdot \text{s}}{(1.16 \times 10^{-21} \text{ kg} \cdot \text{m/s})(4\pi)} = 4.56 \times 10^{-14} \text{ m}.$$

Note that non-relativistic treatment has been used, which is justified because the kinetic energy is only $\frac{(1.6 \times 10^{-13}) \times 100\%}{1.50 \times 10^{-10}} = 0.11\%$ of the rest energy.

- 5-21 (a) The woman tries to hold a pellet within some horizontal region Δx_i and directly above the spot on the floor. The uncertainty principle requires her to give a pellet some x velocity at least as large as $\Delta v_x = \frac{\hbar}{2m\Delta x_i}$. When the pellet hits the floor at time

t , the total miss distance is $\Delta x_{\text{total}} = \Delta x_i + \Delta v_x t = \Delta x_i + \left(\frac{\hbar}{2m\Delta x_i}\right) \sqrt{\frac{2H}{g}}$. The minimum

value of the function Δx_{total} occurs for $\frac{d(\Delta x_{\text{total}})}{d(\Delta x_i)} = 0$ or $1 - \frac{\hbar}{2m} \sqrt{\frac{2H}{g}} (\Delta x_i)^{-2} = 0$.



$$\text{We find } \Delta x_i = \left(\frac{\hbar}{2m}\right)^{1/2} \left(\frac{2H}{g}\right)^{1/4}.$$

- (b) For $H = 2.0$ m, $m = 0.50$ g, $\Delta x_{\text{total}} = 5.2 \times 10^{-16}$ m.

5-23 (a) $\Delta p \Delta x = m \Delta v \Delta x \geq \frac{\hbar}{2}$
 $\Delta v \geq \frac{h}{4\pi m \Delta x} = \frac{2\pi \text{ J} \cdot \text{s}}{4\pi (2 \text{ kg})(1 \text{ m})} = 0.25 \text{ m/s}$

- (b) The duck might move by $(0.25 \text{ m/s})(5 \text{ s}) = 1.25$ m. With original position uncertainty of 1m, we can think of Δx growing to $1 \text{ m} + 1.25 \text{ m} = 2.25$ m.

- 5-25 To find the energy width of the γ -ray use $\Delta E \Delta t \geq \frac{\hbar}{2}$ or

$$\Delta E \geq \frac{\hbar}{2\Delta t} \geq \frac{6.58 \times 10^{-16} \text{ eV} \cdot \text{s}}{(2)(0.10 \times 10^{-9} \text{ s})} \geq 3.29 \times 10^{-6} \text{ eV}.$$

As the intrinsic energy width of $\sim \pm 3 \times 10^{-6}$ eV is so much less than the experimental resolution of ± 5 eV, the intrinsic width can't be measured using this method.

- 5-27 For a single slit with width a , minima are given by $\sin\theta = \frac{n\lambda}{a}$ where $n = 1, 2, 3, \dots$ and $\sin\theta \approx \tan\theta = \frac{x}{L}$, $\frac{x_1}{L} = \frac{\lambda}{a}$ and $\frac{x_2}{L} = \frac{2\lambda}{a} \Rightarrow \frac{x_2 - x_1}{L} = \frac{\lambda}{a}$ or

$$\lambda = \frac{a\Delta x}{L} = \frac{5 \text{ \AA} \times 2.1 \text{ cm}}{20 \text{ cm}} = 0.525 \text{ \AA}$$

$$E = \frac{p^2}{2m} = \frac{h^2}{2m\lambda^2} = \frac{(hc)^2}{2mc^2\lambda^2} = \frac{(1.24 \times 10^4 \text{ eV} \cdot \text{\AA})^2}{2(5.11 \times 10^5 \text{ eV})(0.525 \text{ \AA})^2} = 546 \text{ eV}$$

- 5-29 With *one* slit open $P_1 = |\Psi_1|^2$ or $P_2 = |\Psi_2|^2$. With both slits open, $P = |\Psi_1 + \Psi_2|^2$. At a maximum, the wavefunctions are in phase so

$$P_{\max} = (|\Psi_1| + |\Psi_2|)^2.$$

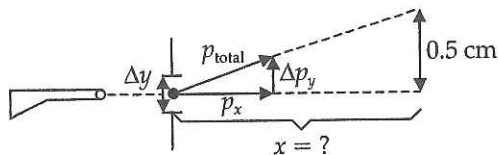
At a minimum, the wavefunctions are out of phase and

$$P_{\min} = (|\Psi_1| - |\Psi_2|)^2.$$

Now $\frac{P_1}{P_2} = \frac{|\Psi_1|^2}{|\Psi_2|^2} = 25$ or $\frac{|\Psi_1|}{|\Psi_2|} = 5$, and

$$\frac{P_{\max}}{P_{\min}} = \frac{(|\Psi_1| + |\Psi_2|)^2}{(|\Psi_1| - |\Psi_2|)^2} = \frac{(5|\Psi_2| + |\Psi_2|)^2}{(5|\Psi_2| - |\Psi_2|)^2} = \frac{6^2}{4^2} = \frac{36}{16} = 2.25.$$

5-31



$\Delta y \Delta p_y \sim \hbar$ $\Delta p_y = \frac{\hbar}{\Delta y}$. From the diagram, because the momentum triangle and space triangle are similar, $\frac{\Delta p_y}{p_x} = \frac{0.5 \text{ cm}}{x}$;

$$x = \frac{(0.5 \text{ cm})p_x}{\Delta p_y} = \frac{(0.5 \text{ cm})p_x \Delta y}{\hbar} = \frac{(0.5 \times 10^{-2} \text{ m})(0.001 \text{ kg})(100 \text{ m/s})(2 \times 10^{-3} \text{ m})}{1.05 \times 10^{-34} \text{ J} \cdot \text{s}}$$

$$= 9.5 \times 10^{27} \text{ m}$$

Once again we see that the uncertainty relation has no observable consequences for macroscopic systems.

5-33 From the uncertainty principle, $\Delta E \Delta t \sim \hbar$ $\Delta mc^2 \Delta t = \hbar$. Therefore,

$$\frac{\Delta m}{m} = \frac{\hbar}{2\pi c^2 \Delta t m} = \frac{\hbar}{2\pi \Delta t E_{\text{rest}}} = \frac{6.63 \times 10^{-34} \text{ J}\cdot\text{s}}{2\pi(8.7 \times 10^{-17} \text{ s})(135 \times 10^6 \text{ eV})(1.6 \times 10^{-19} \text{ J/eV})} = 5.62 \times 10^{-8}$$

5-35 (a) $f(x) = \frac{1}{\sqrt{2\pi}} \int_{-\infty}^{+\infty} a(k) e^{ikx} dk = \frac{A}{\sqrt{2\pi}} \int_{-\infty}^{+\infty} e^{-\alpha^2(k-k_0)^2} e^{ikx} dk = \frac{A}{\sqrt{2\pi}} e^{-\alpha^2 k_0^2} \int_{-\infty}^{+\infty} e^{-\alpha^2(k^2 - (2k_0 + ix/\alpha^2)k)} dk$.

Now complete the square in order to get the integral into the standard form

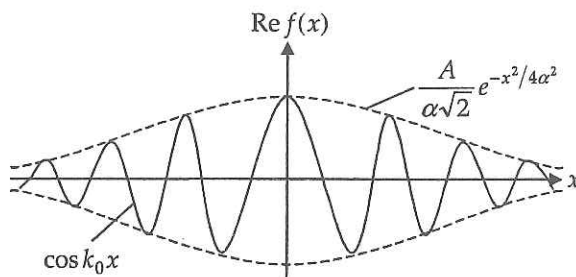
$$\int_{-\infty}^{+\infty} e^{-az^2} dz:$$

$$e^{-\alpha^2(k^2 - (2k_0 + ix/\alpha^2)k)} = e^{+\alpha^2(k_0 + ix/2\alpha^2)^2} e^{-\alpha^2(k - (k_0 + ix/2\alpha^2))^2}$$

$$\begin{aligned} f(x) &= \frac{A}{\sqrt{2\pi}} e^{-\alpha^2 k_0^2} e^{+\alpha^2(k_0 + ix/2\alpha^2)^2} \int_{-\infty}^{+\infty} e^{-\alpha^2(k - (k_0 + ix/2\alpha^2))^2} dk \\ &= \frac{A}{\sqrt{2\pi}} e^{-x^2/4\alpha^2} e^{ik_0 x} \int_{z=-\infty}^{+\infty} e^{-\alpha^2 z^2} dz \end{aligned}$$

where $z = k - \left(k_0 + \frac{ix}{2\alpha^2}\right)$. Since $\int_{z=-\infty}^{+\infty} e^{-\alpha^2 z^2} dz = \frac{\pi^{1/2}}{\alpha}$, $f(x) = \frac{A}{\alpha\sqrt{2}} e^{-x^2/4\alpha^2} e^{ik_0 x}$. The real

part of $f(x)$, $\text{Re } f(x)$ is $\text{Re } f(x) = \frac{A}{\alpha\sqrt{2}} e^{-x^2/4\alpha^2} \cos k_0 x$ and is a gaussian envelope multiplying a harmonic wave with wave number k_0 . A plot of $\text{Re } f(x)$ is shown below:



Comparing $\frac{A}{\alpha\sqrt{2}} e^{-x^2/4\alpha^2}$ to $Ae^{-(x/2\Delta x)^2}$ implies $\Delta x = \alpha$.

(c) By same reasoning because $\alpha^2 = \frac{1}{4\Delta k^2}$, $\Delta k = \frac{1}{2\alpha}$. Finally $\Delta x \Delta k = \alpha \left(\frac{1}{2\alpha}\right) = \frac{1}{2}$.

5-37 We find the speed of each electron from energy conservation in the firing process:

$$0 = K_f + U_f = \frac{1}{2}mv^2 - eV$$

$$v = \sqrt{\frac{2eV}{m}} = \sqrt{\frac{2(1.6 \times 10^{-19} \text{ C})(45 \text{ V})}{9.11 \times 10^{-31} \text{ kg}}} = 3.98 \times 10^6 \text{ m/s}$$

The time of flight is $\Delta t = \frac{\Delta x}{v} = \frac{0.28 \text{ m}}{3.98 \times 10^6 \text{ m/s}} = 7.04 \times 10^{-8} \text{ s}$. The current when electrons are 28 cm apart is $I = \frac{q}{t} = \frac{e}{\Delta t} = \frac{1.6 \times 10^{-19} \text{ C}}{7.04 \times 10^{-8} \text{ s}} = 2.27 \times 10^{-12} \text{ A}$.

6

Quantum Mechanics in One Dimension

- 6-1 (a) Not acceptable – diverges as $x \rightarrow \infty$.
 (b) Acceptable.
 (c) Acceptable.
 (d) Not acceptable – not a single-valued function.
 (e) Not acceptable – the wave is discontinuous (as is the slope).

6-3 (a) $A \sin\left(\frac{2\pi x}{\lambda}\right) = A \sin(5 \times 10^{10} x)$ so $\left(\frac{2\pi}{\lambda}\right) = 5 \times 10^{10} \text{ m}^{-1}$, $\lambda = \frac{2\pi}{5 \times 10^{10}} = 1.26 \times 10^{-10} \text{ m}$.

(b) $p = \frac{h}{\lambda} = \frac{6.626 \times 10^{-34} \text{ Js}}{1.26 \times 10^{-10} \text{ m}} = 5.26 \times 10^{-24} \text{ kg m/s}$

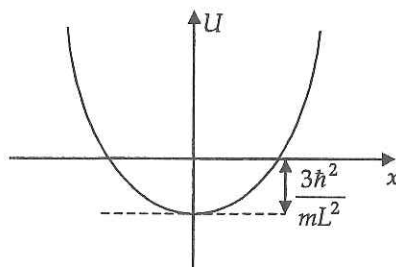
(c) $K = \frac{p^2}{2m}$ $m = 9.11 \times 10^{-31} \text{ kg}$
 $K = \frac{(5.26 \times 10^{-24} \text{ kg m/s})^2}{(2 \times 9.11 \times 10^{-31} \text{ kg})} = 1.52 \times 10^{-17} \text{ J}$
 $K = \frac{1.52 \times 10^{-17} \text{ J}}{1.6 \times 10^{-19} \text{ J/eV}} = 95 \text{ eV}$

- 6-5 (a) Solving the Schrödinger equation for U with $E = 0$ gives

$$U = \left(\frac{\hbar^2}{2m}\right) \left(\frac{d^2\psi}{dx^2}\right) / \psi$$

If $\psi = Ae^{-x^2/L^2}$ then $\frac{d^2\psi}{dx^2} = (4Ax^3 - 6AxL^2) \left(\frac{1}{L^4}\right) e^{-x^2/L^2}$, $U = \left(\frac{\hbar^2}{2mL^2}\right) \left(\frac{4x^2}{L^2} - 6\right)$.

- (b) $U(x)$ is a parabola centered at $x=0$ with $U(0) = \frac{-3\hbar^2}{mL^2} < 0$:



- 6-7 Since the particle is confined to the box, Δx can be no larger than L , the box length. With $n=0$, the particle energy $E_n = \frac{n^2\hbar^2}{8mL^2}$ is also zero. Since the energy is all kinetic, this implies $\langle p_x^2 \rangle = 0$. But $\langle p_x \rangle = 0$ is expected for a particle that spends equal time moving left as right, giving $\Delta p_x = \sqrt{\langle p_x^2 \rangle - \langle p_x \rangle^2} = 0$. Thus, for this case $\Delta p_x \Delta x = 0$, in violation of the uncertainty principle.

6-9 $E_n = \frac{n^2\hbar^2}{8mL^2}$, so $\Delta E = E_2 - E_1 = \frac{3\hbar^2}{8mL^2}$

$$\Delta E = (3) \frac{(1240 \text{ eV nm}/c)^2}{8(938.28 \times 10^6 \text{ eV}/c^2)(10^{-5} \text{ nm})^2} = 6.14 \text{ MeV}$$

$$\lambda = \frac{hc}{\Delta E} = \frac{1240 \text{ eV nm}}{6.14 \times 10^6 \text{ eV}} = 2.02 \times 10^{-4} \text{ nm}$$

This is the gamma ray region of the electromagnetic spectrum.

- 6-11 In the present case, the box is displaced from $(0, L)$ by $\frac{L}{2}$. Accordingly, we may obtain the wavefunctions by replacing x with $x - \frac{L}{2}$ in the wavefunctions of Equation 6.18. Using

$$\sin\left[\left(\frac{n\pi}{L}\right)\left(x - \frac{L}{2}\right)\right] = \sin\left[\left(\frac{n\pi x}{L}\right) - \frac{n\pi}{2}\right] = \sin\left(\frac{n\pi x}{L}\right)\cos\left(\frac{n\pi}{2}\right) - \cos\left(\frac{n\pi x}{L}\right)\sin\left(\frac{n\pi}{2}\right)$$

we get for $-\frac{L}{2} \leq x \leq \frac{L}{2}$

$$\begin{aligned} \psi_1(x) &= \left(\frac{2}{L}\right)^{1/2} \cos\left(\frac{\pi x}{L}\right); & P_1(x) &= \left(\frac{2}{L}\right) \cos^2\left(\frac{\pi x}{L}\right) \\ \psi_2(x) &= \left(\frac{2}{L}\right)^{1/2} \sin\left(\frac{2\pi x}{L}\right); & P_2(x) &= \left(\frac{2}{L}\right) \sin^2\left(\frac{2\pi x}{L}\right) \\ \psi_3(x) &= \left(\frac{2}{L}\right)^{1/2} \cos\left(\frac{3\pi x}{L}\right); & P_3(x) &= \left(\frac{2}{L}\right) \cos^2\left(\frac{3\pi x}{L}\right) \end{aligned}$$

- 6-13 (a) Proton in a box of width $L = 0.200 \text{ nm} = 2 \times 10^{-10} \text{ m}$

$$E_1 = \frac{h^2}{8m_p L^2} = \frac{(6.626 \times 10^{-34} \text{ J}\cdot\text{s})^2}{8(1.67 \times 10^{-27} \text{ kg})(2 \times 10^{-10} \text{ m})^2} = 8.22 \times 10^{-22} \text{ J}$$

$$= \frac{8.22 \times 10^{-22} \text{ J}}{1.60 \times 10^{-19} \text{ J/eV}} = 5.13 \times 10^{-3} \text{ eV}$$

- (b) Electron in the same box:

$$E_1 = \frac{h^2}{8m_e L^2} = \frac{(6.626 \times 10^{-34} \text{ J}\cdot\text{s})^2}{8(9.11 \times 10^{-31} \text{ kg})(2 \times 10^{-10} \text{ m})^2} = 1.506 \times 10^{-18} \text{ J} = 9.40 \text{ eV}.$$

- (c) The electron has a much higher energy because it is much less massive.

6-15 (a) $U = \left(\frac{e^2}{4\pi\epsilon_0 d} \right) \left[-1 + \frac{1}{2} - \frac{1}{3} + \left(-1 + \frac{1}{2} \right) + (-1) \right] = \frac{(-7/3)e^2}{4\pi\epsilon_0 d} = \frac{(-7/3)ke^2}{d}$

(b) $K = 2E_1 = \frac{2h^2}{8m \times 9d^2} = \frac{h^2}{36md^2}$

(c) $E = U + K$ and $\frac{dE}{dd} = 0$ for a minimum $\left[\frac{(+7/3)e^2 k}{d^2} \right] - \frac{h^2}{18md^3} = 0$

$$d = \frac{3h^2}{(7)(18ke^2 m)} \text{ or } d = \frac{h^2}{42mke^2}$$

$$d = \frac{(6.63 \times 10^{-34} \text{ J}\cdot\text{s})^2}{(42)(9.11 \times 10^{-31} \text{ kg})(9 \times 10^9 \text{ N}\cdot\text{m}^2 \cdot \text{C}^{-2})(1.6 \times 10^{-19} \text{ C})^2} = 0.5 \times 10^{-10} \text{ m} = 0.050 \text{ nm}$$

- (d) Since the lithium spacing is a , where $Na^3 = V$ and the density is $\frac{Nm}{V}$ where m is the mass of one atom, we get

$$a = \left(\frac{Vm}{Nm} \right)^{1/3} = \left(\frac{m}{\text{density}} \right)^{1/3} = \left(1.66 \times 10^{-27} \text{ kg} \times \frac{7}{530 \text{ kg/m}^3} \right)^{1/3} = 2.8 \times 10^{-10} \text{ m}$$

$$= 0.28 \text{ nm}$$

(2.8 times larger than $2d$)

- 6-17 (a) The wavefunctions and probability densities are the same as those shown in the two lower curves in Figure 6.16 of the text.

(b) $P_1 = \int_{1.5 \text{ \AA}}^{3.5 \text{ \AA}} |\psi|^2 dx = \frac{2}{10 \text{ \AA}} \int_{1.5 \text{ \AA}}^{3.5 \text{ \AA}} \sin^2 \left(\frac{\pi x}{10} \right) dx$

$$\frac{1}{5} \left[\frac{x}{2} - \frac{10}{4\pi} \sin \left(\frac{\pi x}{5} \right) \right]_{1.5}^{3.5}$$

In the above result we used $\int \sin^2 ax dx = \frac{x}{2} - \frac{1}{4a} \sin(2ax)$. Therefore,

$$P_1 = \frac{1}{10} \left[x - \frac{5}{\pi} \sin\left(\frac{\pi x}{5}\right) \right]_{1.5}^{3.5} = \frac{1}{10} \left\{ 3.5 - \frac{5}{\pi} \sin\left[\frac{\pi(3.5)}{5}\right] - 1.5 + \frac{5}{\pi} \sin\left[\frac{\pi(1.5)}{5}\right] \right\}$$

$$= \frac{1}{10} \left[2.0 + \frac{5}{\pi} (\sin 0.3\pi - \sin 0.7\pi) \right] = \frac{1}{10} [2.00 + 0.0] = 0.200$$

$$(c) \quad P_2 = \frac{1}{5} \int_{1.5}^{3.5} \sin^2\left(\frac{\pi x}{5}\right) dx = \frac{1}{5} \left[\frac{x}{2} - \frac{5}{4\pi} \sin(0.4\pi x) \right]_{1.5}^{3.5} = \frac{1}{10} \left[x - \frac{5}{2\pi} \sin(0.4\pi x) \right]_{1.5}^{3.5}$$

$$= \frac{1}{10} \{ 2.0 + (0.798) [\sin[0.4\pi(1.5)] - \sin[0.4\pi(3.5)]] \} = 0.351$$

$$(d) \quad \text{Using } E = \frac{n^2 h^2}{8mL^2} \text{ we find } E_1 = 0.377 \text{ eV and } E_2 = 1.51 \text{ eV.}$$

6-19 The allowed energies for this system are given by Equation 6.17, or $E_n = \frac{n^2 \pi^2 \hbar^2}{2mL^2} = \frac{n^2 h^2}{8mL^2}$. Using $E_n = 10^{-3} \text{ J}$, $m = 10^{-3} \text{ kg}$, $L = 10^{-2} \text{ m}$ and solving for n gives

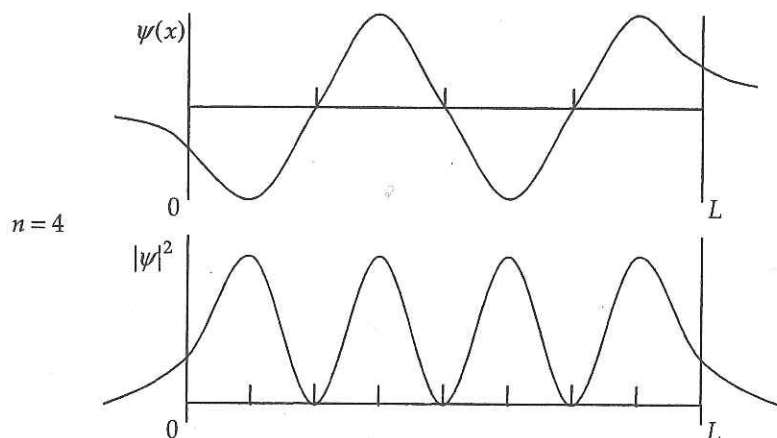
$$n = \frac{\{ 8(10^{-3} \text{ kg})(10^{-2} \text{ m})^2 (10^{-3} \text{ J}) \}^{1/2}}{6.63 \times 10^{-34} \text{ J}\cdot\text{s}} = 4.27 \times 10^{28}.$$

The excitation energy is $\Delta E = E_{n+1} - E_n$, or

$$\Delta E = \frac{h^2}{8mL^2} \{ (n+1)^2 - n^2 \} = \left(\frac{h^2}{8mL^2} \right) \{ 2n+1 \} = E_n \left(\frac{2n+1}{n^2} \right) \approx \frac{2}{n} E_n \text{ for } n \gg 1.$$

$$\text{Thus, } \Delta E \approx \frac{(2)(10^{-3} \text{ J})}{4.27 \times 10^{28}} = 4.69 \times 10^{-32} \text{ J.}$$

6-21 $n = 4$



Note that the $n = 4$ wavefunction has three nodes and is antisymmetric about the midpoint of the well.

- 6-23 Inside the well, the particle is free and the Schrödinger waveform is trigonometric with wavenumber $k = \left(\frac{2mE}{\hbar^2}\right)^{1/2}$:

$$\psi(x) = A \sin kx + B \cos kx \quad 0 \leq x \leq L.$$

The infinite wall at $x=0$ requires $\psi(0) = B = 0$. Beyond $x=L$, $U(x) = U$ and the Schrödinger equation $\frac{d^2\psi}{dx^2} = \left(\frac{2m}{\hbar^2}\right)\{U - E\}\psi(x)$, which has exponential solutions for $E < U$

$$\psi(x) = Ce^{-\alpha x} + De^{+\alpha x}, \quad x > L$$

where $\alpha = \left[\frac{2m(U-E)}{\hbar^2}\right]^{1/2}$. To keep ψ bounded at $x = \infty$ we must take $D = 0$. At $x = L$, continuity of ψ and $\frac{d\psi}{dx}$ demands

$$\begin{aligned} A \sin kL &= Ce^{-\alpha L} \\ kA \cos kL &= -\alpha Ce^{-\alpha L} \end{aligned}$$

Dividing one by the other gives an equation for the allowed particle energies: $k \cot kL = -\alpha$.

The dependence on E (or k) is made more explicit by noting that $k^2 + \alpha^2 = \frac{2mU}{\hbar^2}$, which

allows the energy condition to be written $k \cot kL = -\left[\left(\frac{2mU}{\hbar^2}\right) - k^2\right]^{1/2}$. Multiplying by L ,

squaring the result, and using $\cot^2 \theta + 1 = \csc^2 \theta$ gives $(kL)^2 \csc^2(kL) = \frac{2mUL^2}{\hbar^2}$ from which we

obtain $\frac{kL}{\sin kL} = \left(\frac{2mUL^2}{\hbar^2}\right)^{1/2}$. Since $\frac{\theta}{\sin \theta}$ is never smaller than unity for any value of θ , there

can be no bound state energies if $\frac{2mUL^2}{\hbar^2} < 1$.

- 6-25 At its limits of vibration $x = \pm A$ the classical oscillator has all its energy in potential form:

$E = \frac{1}{2}m\omega^2 A^2$ or $A = \left(\frac{2E}{m\omega^2}\right)^{1/2}$. If the energy is quantized as $E_n = \left(n + \frac{1}{2}\right)\hbar\omega$, then the

corresponding amplitudes are $A_n = \left[\frac{(2n+1)\hbar}{m\omega}\right]^{1/2}$.

- 6-29 (a) Normalization requires $1 = \int_{-\infty}^{\infty} |\psi|^2 dx = C^2 \int_0^{\infty} e^{-2x} (1 - e^{-x})^2 dx = C^2 \int_0^{\infty} (e^{-2x} - 2e^{-3x} + e^{-4x}) dx$.

The integrals are elementary and give $1 = C^2 \left\{ \frac{1}{2} - 2\left(\frac{1}{3}\right) + \frac{1}{4} \right\} = \frac{C^2}{12}$. The proper units for C are those of $(\text{length})^{-1/2}$ thus, normalization requires $C = (12)^{1/2} \text{ nm}^{-1/2}$.

- (b) The most likely place for the electron is where the probability $|\psi|^2$ is largest. This is also where ψ itself is largest, and is found by setting the derivative $\frac{d\psi}{dx}$ equal zero:

$$0 = \frac{d\psi}{dx} = C\{-e^{-x} + 2e^{-2x}\} = Ce^{-x}\{2e^{-x} - 1\}.$$

The RHS vanishes when $x = \infty$ (a minimum), and when $2e^{-x} = 1$, or $x = \ln 2$ nm. Thus, the most likely position is at $x_p = \ln 2$ nm = 0.693 nm.

- (c) The average position is calculated from

$$\langle x \rangle = \int_{-\infty}^{\infty} x|\psi|^2 dx = C^2 \int_0^{\infty} xe^{-2x}(1 - e^{-x})^2 dx = C^2 \int_0^{\infty} x(e^{-2x} - 2e^{-3x} + e^{-4x}) dx.$$

The integrals are readily evaluated with the help of the formula $\int_0^{\infty} xe^{-ax} dx = \frac{1}{a^2}$ to get

$$\langle x \rangle = C^2 \left\{ \frac{1}{4} - 2\left(\frac{1}{9}\right) + \frac{1}{16} \right\} = C^2 \left\{ \frac{13}{144} \right\}. \text{ Substituting } C^2 = 12 \text{ nm}^{-1} \text{ gives}$$

$$\langle x \rangle = \frac{13}{12} \text{ nm} = 1.083 \text{ nm}.$$

We see that $\langle x \rangle$ is somewhat greater than the most probable position, since the probability density is skewed in such a way that values of x larger than x_p are weighted more heavily in the calculation of the average.

- 6-31 The symmetry of $|\psi(x)|^2$ about $x=0$ can be exploited effectively in the calculation of average values. To find $\langle x \rangle$

$$\langle x \rangle = \int_{-\infty}^{\infty} x|\psi(x)|^2 dx$$

We notice that the integrand is antisymmetric about $x=0$ due to the extra factor of x (an odd function). Thus, the contribution from the two half-axes $x > 0$ and $x < 0$ cancel exactly, leaving $\langle x \rangle = 0$. For the calculation of $\langle x^2 \rangle$, however, the integrand is symmetric and the half-axes contribute equally to the value of the integral, giving

$$\langle x \rangle = \int_0^{\infty} x^2 |\psi|^2 dx = 2C^2 \int_0^{\infty} x^2 e^{-2x/x_0} dx.$$

Two integrations by parts show the value of the integral to be $2\left(\frac{x_0}{2}\right)^3$. Upon substituting for C^2 , we get $\langle x^2 \rangle = 2\left(\frac{1}{x_0}\right)\left(2\right)\left(\frac{x_0}{2}\right)^3 = \frac{x_0^2}{2}$ and $\Delta x = (\langle x^2 \rangle - \langle x \rangle^2)^{1/2} = \left(\frac{x_0^2}{2}\right)^{1/2} = \frac{x_0}{\sqrt{2}}$. In calculating the probability for the interval $-\Delta x$ to $+\Delta x$ we appeal to symmetry once again to write

$$P = \int_{-\Delta x}^{+\Delta x} |\psi|^2 dx = 2C^2 \int_0^{\Delta x} e^{-2x/x_0} dx = -2C^2 \left(\frac{x_0}{2}\right) e^{-2x/x_0} \Big|_0^{\Delta x} = 1 - e^{-\sqrt{2}} = 0.757$$

or about 75.7% independent of x_0 .

- 6-33 (a) Since there is no preference for motion in the leftward sense vs. the rightward sense, a particle would spend equal time moving left as moving right, suggesting $\langle p_x \rangle = 0$.
- (b) To find $\langle p_x^2 \rangle$ we express the average energy as the sum of its kinetic and potential energy contributions: $\langle E \rangle = \left\langle \frac{p_x^2}{2m} \right\rangle + \langle U \rangle = \frac{\langle p_x^2 \rangle}{2m} + \langle U \rangle$. But energy is sharp in the oscillator ground state, so that $\langle E \rangle = E_0 = \frac{1}{2} \hbar \omega$. Furthermore, remembering that $U(x) = \frac{1}{2} m \omega^2 x^2$ for the quantum oscillator, and using $\langle x^2 \rangle = \frac{\hbar}{2m\omega}$ from Problem 6-32, gives $\langle U \rangle = \frac{1}{2} m \omega^2 \langle x^2 \rangle = \frac{1}{4} \hbar \omega$. Then $\langle p_x^2 \rangle = 2m(E_0 - \langle U \rangle) = 2m\left(\frac{\hbar \omega}{4}\right) = \frac{m \hbar \omega}{2}$.
- (c) $\Delta p_x = (\langle p_x^2 \rangle - \langle p_x \rangle^2)^{1/2} = \left(\frac{m \hbar \omega}{2}\right)^{1/2}$

6-35 Applying the momentum operator $[p_x] = \left(\frac{\hbar}{i}\right) \frac{d}{dx}$ to each of the candidate functions yields

- (a) $[p_x]\{A \sin(kx)\} = \left(\frac{\hbar}{i}\right) k \{A \cos(kx)\}$
- (b) $[p_x]\{A \sin(kx) - A \cos(kx)\} = \left(\frac{\hbar}{i}\right) k \{A \cos(kx) + A \sin(kx)\}$
- (c) $[p_x]\{A \cos(kx) + iA \sin(kx)\} = \left(\frac{\hbar}{i}\right) k \{-A \sin(kx) + iA \cos(kx)\}$
- (d) $[p_x]\{e^{ik(x-a)}\} = \left(\frac{\hbar}{i}\right) ik \{e^{ik(x-a)}\}$

In case (c), the result is a multiple of the original function, since

$$-A \sin(kx) + iA \cos(kx) = i\{A \cos(kx) + iA \sin(kx)\}.$$

The multiple is $\left(\frac{\hbar}{i}\right)(ik) = \hbar k$ and is the eigenvalue. Likewise for (d), the operation $[p_x]$ returns the original function with the multiplier $\hbar k$. Thus, (c) and (d) are eigenfunctions of $[p_x]$ with eigenvalue $\hbar k$, whereas (a) and (b) are not eigenfunctions of this operator.

6-37. (a) Normalization requires

$$\begin{aligned} 1 &= \int_{-\infty}^{\infty} |\Psi|^2 dx = C^2 \int_{-\infty}^{\infty} \{\psi_1^* + \psi_2^*\} \{\psi_1 + \psi_2\} dx \\ &= C^2 \left\{ \int |\psi_1|^2 dx + \int |\psi_2|^2 dx + \int \psi_2^* \psi_1 dx + \int \psi_1^* \psi_2 dx \right\} \end{aligned}$$

The first two integrals on the right are unity, while the last two are, in fact, the same integral since ψ_1 and ψ_2 are both real. Using the waveforms for the infinite square well, we find

$$\int \psi_2 \psi_1 dx = \frac{2}{L} \int_0^L \sin\left(\frac{\pi x}{L}\right) \sin\left(\frac{2\pi x}{L}\right) dx = \frac{1}{L} \int_0^L \left\{ \cos\left(\frac{\pi x}{L}\right) - \cos\left(\frac{3\pi x}{L}\right) \right\} dx$$

where, in writing the last line, we have used the trigonometric exponential identities of sine and cosine. Both of the integrals remaining are readily evaluated, and are zero.

Thus, $1 = C^2 \{1 + 0 + 0 + 0\} = 2C^2$, or $C = \frac{1}{\sqrt{2}}$. Since $\psi_{1,2}$ are stationary states, they

develop in time according to their respective energies $E_{1,2}$ as $e^{-iE_t/\hbar}$. Then $\Psi(x, t) = C \{ \psi_1 e^{-iE_1 t/\hbar} + \psi_2 e^{-iE_2 t/\hbar} \}$.

(c) $\Psi(x, t)$ is a stationary state only if it is an eigenfunction of the energy operator $[E] = i\hbar \frac{\partial}{\partial t}$. Applying $[E]$ to Ψ gives

$$[E]\Psi = C \left\{ i\hbar \left(\frac{-iE_1}{\hbar} \right) \psi_1 e^{-iE_1 t/\hbar} + i\hbar \left(\frac{-iE_2}{\hbar} \right) \psi_2 e^{-iE_2 t/\hbar} \right\} = C \{ E_1 \psi_1 e^{-iE_1 t/\hbar} + E_2 \psi_2 e^{-iE_2 t/\hbar} \}.$$

Since $E_1 \neq E_2$, the operations $[E]$ does *not* return a multiple of the wavefunction, and so Ψ is not a stationary state. Nonetheless, we may calculate the average energy for this state as

$$\begin{aligned} \langle E \rangle &= \int \Psi^* [E]\Psi dx = C^2 \int \{ \psi_1^* e^{+iE_1 t/\hbar} + \psi_2^* e^{+iE_2 t/\hbar} \} \{ E_1 \psi_1 e^{-iE_1 t/\hbar} + E_2 \psi_2 e^{-iE_2 t/\hbar} \} dx \\ &= C^2 \{ E_1 \int |\psi_1|^2 dx + E_2 \int |\psi_2|^2 dx \} \end{aligned}$$

with the cross terms vanishing as in part (a). Since $\psi_{1,2}$ are normalized and $C^2 = \frac{1}{2}$

we get finally $\langle E \rangle = \frac{E_1 + E_2}{2}$.

7

Tunneling Phenomena

7-1 (a) The reflection coefficient is the ratio of the reflected intensity to the incident wave intensity, or $R = \frac{|(1/2)(1-i)|^2}{|(1/2)(1+i)|^2}$. But $|1-i|^2 = (1-i)(1-i)^* = (1-i)(1+i) = |1+i|^2 = 2$, so that $R = 1$ in this case.

(b) To the left of the step the particle is free. The solutions to Schrödinger's equation are $e^{\pm ikx}$ with wavenumber $k = \left(\frac{2mE}{\hbar^2}\right)^{1/2}$. To the right of the step $U(x) = U$ and the equation is $\frac{d^2\psi}{dx^2} = \frac{2m}{\hbar^2}(U-E)\psi(x)$. With $\psi(x) = e^{-kx}$, we find $\frac{d^2\psi}{dx^2} = k^2\psi(x)$, so that $k = \left[\frac{2m(U-E)}{\hbar^2}\right]^{1/2}$. Substituting $k = \left(\frac{2mE}{\hbar^2}\right)^{1/2}$ shows that $\left[\frac{E}{(U-E)}\right]^{1/2} = 1$ or $\frac{E}{U} = \frac{1}{2}$.

(c) For 10 MeV protons, $E = 10$ MeV and $m = \frac{938.28 \text{ MeV}}{c^2}$. Using $\hbar = 197.3 \text{ MeV fm}/c$ ($1 \text{ fm} = 10^{-15} \text{ m}$), we find $\delta = \frac{1}{k} = \frac{\hbar}{(2mE)^{1/2}} = \frac{197.3 \text{ MeV fm}/c}{[(2)(938.28 \text{ MeV}/c^2)(10 \text{ MeV})]^{1/2}} = 1.44 \text{ fm}$.

7-3 With $E = 25$ MeV and $U = 20$ MeV, the ratio of wavenumber is

$$\frac{k_1}{k_2} = \left(\frac{E}{E-U}\right)^{1/2} = \left(\frac{25}{25-20}\right)^{1/2} = \sqrt{5} = 2.236. \text{ Then from Problem 7-2 } R = \frac{(\sqrt{5}-1)^2}{(\sqrt{5}+1)^2} = 0.146 \text{ and}$$

$T = 1 - R = 0.854$. Thus, 14.6% of the incoming particles would be reflected and 85.4% would be transmitted. For electrons with the same energy, the transparency and reflectivity of the step are unchanged.

7-5 (a) The transmission probability according to Equation 7.9 is

$$\frac{1}{T(E)} = 1 + \left[\frac{U^2}{4E(U-E)}\right] \sinh^2 \alpha L \text{ with } \alpha = \frac{[2m(U-E)]^{1/2}}{\hbar}. \text{ For } E \ll U, \text{ we find}$$

$$(\alpha L)^2 \approx \frac{2mUL^2}{\hbar^2} \gg 1 \text{ by hypothesis. Thus, we may write } \sinh \alpha L \approx \frac{1}{2}e^{\alpha L}. \text{ Also}$$

$$U - E \approx U, \text{ giving } \frac{1}{T(E)} \approx 1 + \left(\frac{U}{16E}\right)e^{2\alpha L} \approx \left(\frac{U}{16E}\right)e^{2\alpha L} \text{ and a probability for transmission}$$

$$P = T(E) = \left(\frac{16E}{U}\right)e^{-2\alpha L}.$$

- (b) Numerical Estimates: ($\hbar = 1.055 \times 10^{-34}$ Js)
- 1) For $m = 9.11 \times 10^{-31}$ kg, $U - E = 1.60 \times 10^{-21}$ J, $L = 10^{-10}$ m;
 $\alpha = \frac{[2m(U - E)]^{1/2}}{\hbar} = 5.12 \times 10^8 \text{ m}^{-1}$ and $e^{-2\alpha L} = 0.90$
 - 2) For $m = 9.11 \times 10^{-31}$ kg, $U - E = 1.60 \times 10^{-19}$ J, $L = 10^{-10}$ m; $\alpha = 5.12 \times 10^9 \text{ m}^{-1}$
 and $e^{-2\alpha L} = 0.36$
 - 3) For $m = 6.7 \times 10^{-27}$ kg, $U - E = 1.60 \times 10^{-13}$ J, $L = 10^{-15}$ m; $\alpha = 4.4 \times 10^{14} \text{ m}^{-1}$
 and $e^{-2\alpha L} = 0.41$
 - 4) For $m = 8$ kg, $U - E = 1$ J, $L = 0.02$ m; $\alpha = 3.8 \times 10^{34} \text{ m}^{-1}$ and
 $e^{-2\alpha L} = e^{-1.5 \times 10^{33}} \approx 0$

7-7 The continuity requirements from Equation 7.8 are

$$\begin{array}{ll} A + B = C + D & \text{[continuity of } \Psi \text{ at } x = 0\text{]} \\ ikA - ikB = \alpha D - \alpha C & \text{[continuity of } \frac{\partial \Psi}{\partial x} \text{ at } x = 0\text{]} \\ Ce^{-\alpha L} + De^{+\alpha L} = Fe^{ikL} & \text{[continuity of } \Psi \text{ at } x = L\text{]} \\ \alpha De^{+\alpha L} - \alpha Ce^{-\alpha L} = ikFe^{ikL} & \text{[continuity of } \frac{\partial \Psi}{\partial x} \text{ at } x = L\text{]} \end{array}$$

To isolate the transmission amplitude $\frac{F}{A}$, we must eliminate from these relations the unwanted coefficients B , C , and D . Dividing the second line by ik and adding to the first eliminates B , leaving A in terms of C and D . In the same way, dividing the fourth line by α and adding the result to the third line gives D (in terms of F), while subtracting the result from the third line gives C (in terms of F). Combining these results finally yields A :

$$A = \frac{1}{4} Fe^{ikL} \left\{ \left[2 - \left(\frac{\alpha}{ik} + \frac{ik}{\alpha} \right) \right] e^{+\alpha L} + \left[2 + \left(\frac{\alpha}{ik} + \frac{ik}{\alpha} \right) \right] e^{-\alpha L} \right\}. \text{ The transmission probability is } T = \left| \frac{F}{A} \right|^2.$$

Making use of the identities $e^{\pm\alpha L} = \cosh \alpha L \pm \sinh \alpha L$ and $\cosh^2 \alpha L = 1 + \sinh^2 \alpha L$, we obtain

$$\begin{aligned} \frac{1}{T} &= \left| \frac{A}{F} \right|^2 = \frac{1}{4} \left| 2 \cosh \alpha L + i \left(\frac{\alpha}{k} - \frac{k}{\alpha} \right) \sinh \alpha L \right|^2 = \cosh^2 \alpha L + \frac{1}{4} \left(\frac{\alpha}{k} - \frac{k}{\alpha} \right)^2 \sinh^2 \alpha L \\ &= 1 + \frac{1}{4} \left[\frac{U - E}{E} + \frac{E}{U - E} + 2 \right] \sinh^2 \alpha L = 1 + \frac{1}{4} \left[\frac{U^2}{E(U - E)} \right] \sinh^2 \alpha L \end{aligned}$$

- 7-11 (a) The matter wave reflected from the trailing edge of the well ($x = L$) must travel the extra distance $2L$ before combining with the wave reflected from the leading edge ($x = 0$). For $\lambda_2 = 2L$, these two waves interfere destructively since the latter suffers a phase shift of 180° upon reflection, as discussed in Example 7.3.
- (b) The wave functions in all three regions are free particle plane waves. In regions 1 and 3 where $U(x) = U$ we have

$$\begin{aligned} \Psi(x, t) &= Ae^{i(k'x - \omega t)} + Be^{i(-k'x - \omega t)} & x < 0 \\ \Psi(x, t) &= Fe^{i(k'x - \omega t)} + Ge^{i(-k'x - \omega t)} & x < 0 \end{aligned}$$

with $k' = \frac{[2m(E-U)]^{1/2}}{\hbar}$. In this case $G = 0$ since the particle is incident from the left. In region 2 where $U(x) = 0$ we have

$$\Psi(x, t) = Ce^{i(-kx - \omega t)} + De^{i(kx - \omega t)} \quad 0 < x < L$$

with $k = \frac{(2mE)^{1/2}}{\hbar} = \frac{2\pi}{\lambda_2} = \frac{\pi}{L}$ for the case of interest. The wave function and its slope are continuous everywhere, and in particular at the well edges $x = 0$ and $x = L$. Thus, we must require

$$\begin{array}{ll} A + B = C + D & \left[\text{continuity of } \Psi \text{ at } x = 0 \right] \\ k'A - k'B = kD - kC & \left[\text{continuity of } \frac{\partial \Psi}{\partial x} \text{ at } x = 0 \right] \\ Ce^{-ikL} + De^{ikL} = Fe^{ik'L} & \left[\text{continuity of } \Psi \text{ at } x = L \right] \\ kDe^{ikL} - kCe^{-ikL} = k'Fe^{ik'L} & \left[\text{continuity of } \frac{\partial \Psi}{\partial x} \text{ at } x = L \right] \end{array}$$

For $kL = \pi$, $e^{\pm ikL} = -1$ and the last two requirements can be combined to give $kD - kC = k'C + k'D$. Substituting this into the second requirement implies $A - B = C + D$, which is consistent with the first requirement only if $B = 0$, i.e., no reflected wave in region 1.

7-13 As in Problem 7-12, waveform continuity and the slope condition at the site of the delta well demand $A + B = F$ and $ik(A - B) - ikF = -\left(\frac{2mS}{\hbar^2}\right)F$. Dividing the second of these equations by ik and subtracting from the first gives $2B + F = F + \frac{(2mS/\hbar^2)F}{ik}$, or $B = -i\left(\frac{mS}{\hbar^2 k}\right)F = -iF\left(\frac{-E_0}{E}\right)^{1/2}$.

Thus, the reflection coefficient R is $R(E) = \frac{|B|^2}{|A|^2} = \frac{|B|^2}{|F|^2} \frac{|F|^2}{|A|^2} = \left(\frac{-E_0}{E}\right) \left[1 + \left(\frac{-E_0}{E}\right)\right]^{-1}$. Then, with $T(E)$ from Problem 7-12, $T(E) = \left[1 + \left(\frac{-E_0}{E}\right)\right]^{-1}$, we find $R(E) + T(E) = \left(1 - \frac{E_0}{E}\right) \left[1 + \left(\frac{-E_0}{E}\right)\right]^{-1} = 1$.

7-15 Divide the barrier region into N subintervals of length $\Delta x = x_{i+1} - x_i$. For the barrier in the i^{th} subinterval, denote by A_i and F_i the incident and transmitted wave amplitudes, respectively.

The transmission coefficient for this interval is then $T_i = \frac{|F_i|^2}{|A_i|^2}$, and that for the entire barrier is

$$T(E) = \frac{|F_N|^2}{|A_1|^2}. \text{ Now consider the product } \prod T_i = T_1 T_2 T_3 \dots T_N = \left(\frac{|F_1|^2}{|A_1|^2}\right) \left(\frac{|F_2|^2}{|A_2|^2}\right) \left(\frac{|F_3|^2}{|A_3|^2}\right) \dots \left(\frac{|F_N|^2}{|A_N|^2}\right).$$

Assuming the transmitted wave intensity for one barrier becomes the incident wave intensity for the next, we have $|F_1|^2 = |A_2|^2$, $|F_2|^2 = |A_3|^2$ etc., so that $T(E) = \frac{|F_N|^2}{|A_1|^2} = T_1 T_2 T_3 \dots T_N$. Next, we

assume that Δx is sufficiently small and that $U(x)$ is sensibly constant over each interval (so that the square barrier result can be used for T_i), yet large enough to approximate $\sinh \alpha_i \Delta x$ with $\frac{1}{2}e^{\alpha_i \Delta x}$, where α_i is the value taken by α in the i^{th} subinterval: $\alpha_i = \frac{[2m(U_i - E)]^{1/2}}{\hbar}$.

Then, $\frac{1}{T_i} = 1 + \left[\frac{U_i^2}{4E(U_i - E)} \right] \sinh^2(\alpha_i \Delta x) \approx \left[\frac{U_i^2}{16E(U_i - E)} \right] e^{2\alpha_i \Delta x}$ and the transmission coefficient for the entire barrier becomes $T(E) \approx \Pi \left\{ \left[\frac{16E(U_i - E)}{U_i^2} \right] e^{-2\alpha_i \Delta x} \right\} \approx \left[\frac{\Pi 16E(U_i - E)}{U_i^2} \right] e^{-\Sigma 2\alpha_i \Delta x}$. To recover Equation 7.10, we approximate the sum in the exponential by an integral, and note that the product in square brackets is a term of order 1: $T(E) \sim e^{-\Sigma 2\alpha_i \Delta x} \approx e^{-\int 2\alpha(x) dx}$ where now $\alpha(x) = \frac{2m[U(x) - E]^{1/2}}{\hbar}$.

7-17 The collision frequency f is the reciprocal of the transit time for the alpha particle crossing the nucleus, or $f = \frac{v}{2R}$, where v is the speed of the alpha. Now v is found from the kinetic energy which, inside the nucleus, is not the total energy E but the difference $E - U$ between the total energy and the potential energy representing the bottom of the nuclear well. At the nuclear radius $R = 9$ fm, the Coulomb energy is

$$\frac{k(Ze)(2e)}{R} = 2Z \left(\frac{ke^2}{a_0} \right) \left(\frac{a_0}{R} \right) = 2(88)(27.2 \text{ eV}) \left(\frac{5.29 \times 10^{-4} \text{ fm}}{9 \text{ fm}} \right) = 28.14 \text{ MeV}.$$

From this we conclude that $U = -1.86$ MeV to give a nuclear barrier of 30 MeV overall. Thus an alpha with $E = 4.05$ MeV has kinetic energy $4.05 + 1.86 = 5.91$ MeV inside the nucleus. Since the alpha particle has the combined mass of 2 protons and 2 neutrons, or about $3755.8 \text{ MeV}/c^2$ this kinetic energy represents a speed

$$v = \left(\frac{2E_k}{m} \right)^{1/2} = \left[\frac{2(5.91)}{3755.8 \text{ MeV}/c^2} \right]^{1/2} = 0.056c.$$

Thus, we find for the collision frequency $f = \frac{v}{2R} = \frac{0.056c}{2(9 \text{ fm})} = 9.35 \times 10^{20} \text{ Hz}$.